

# OSIRIS Collision Package Overview

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# Why collisions are needed

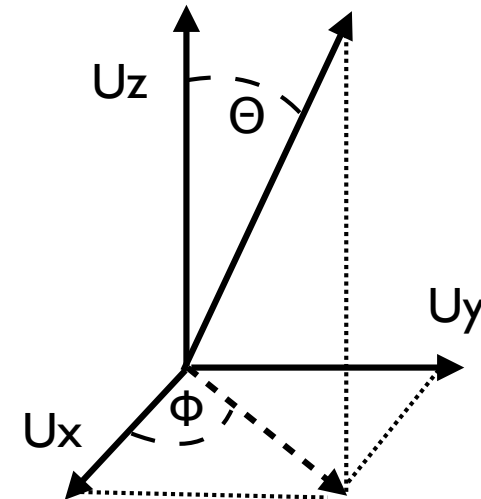


- PIC makes simulating plasmas computationally possible by restricting field to a grid and using finite-size macroparticles.
- Effects smaller than a cell size (rarely smaller than a Debye length) can't be resolved.
- Pure PIC has no concept of real number density; one simulation represents a range of physical systems, up to a scaling factor.
- Discrete particles effects are simulated using Monte Carlo (stochastic) methods, assuming the information which was discarded was essentially random
- (Note PIC \*DOES\* have collisions, but they're distinct in scale and behavior from physical collisions.)

# Monte Carlo Binary Coulomb Collision Model



- Particles are randomly paired, and collided off each other using fluid and particle data
- In center-of-momentum frame, collision is a rotation in p-space
- $\Theta$  is a random variable with variance given by collision frequency (1);  $\Phi$  is random and uniform
- By randomly sampling many pairs, approximates an integral over the distribution function, and so solves the collision term in the Landau form (2).
- Assumes  $\Theta \ll 1$



$$(1) \quad \langle \tan(\Theta) \rangle^2 = \frac{2\pi e_a^2 e_b^2 n \ln(\lambda)}{\mu_{ab}^2 u^3} \Delta t$$

$$(2) \quad \left( \frac{\partial f}{\partial t} \right)_{\text{col}} = - \sum_{\beta} \frac{\partial}{\partial v_j} \frac{2\pi e_a^2 e_b^2 \ln(\lambda)}{m_a} \int d\mathbf{v}' \left[ \frac{\delta_{jk}}{u} - \frac{u_j u_k}{u^3} \right] \left[ \frac{f_a}{m_b} \frac{\partial f_b(\mathbf{v}')}{\partial v'_k} - \frac{f_b(\mathbf{v}')}{m_a} \frac{\partial f_a}{\partial v_k} \right]$$

# Nanbu extension to large collision events



- Collision model of Nanbu
- Question is how to model many small-angle scatterings from Coulomb Collisions in a single simulated collision timestep
- Ansatz assumption that the distribution of the scattering angle is given by equation 1
- Equations 3, 4 then follow analytically; though 4 is non-invertible, various approximations in different regimes can be used to efficiently numerically solve these

$$(1) \quad \left\langle \sin^2 \frac{\chi_N}{2} \right\rangle = \frac{1}{2} (1 - e^{-s})$$

$$(2) \quad s = 4\pi \ln \Lambda \left( \frac{q_1 q_2}{\mu_{12}} \right)^2 n_2 v_{rel}^{-3} \Delta t \quad (s \sim \nu_{coll} \Delta t)$$

$$(3) \quad \cos \chi = \frac{1}{A} \ln (e^{-A} + 2U \sinh A)$$

$$(4) \quad \coth A - A^{-1} = e^{-s}$$

# Pérez Relativistic and Low Temperature Corrections



- Equation 2 is not however relativistically correct (no distinction is made between  $p$  and  $mv$ , for instance)
- Pérez et al. solved this by using a relativistic invariant to arrive at equation 5
- Working in momentum space the collision remains a rotation, and so equations 3, 4 remain valid with the new  $s$

$$(5) \quad s_{12} = \frac{4\pi n_2 \Delta t \ln \Lambda (q_1 q_2)^2}{c^4 m_1 \gamma_1 m_2 \gamma_2} \frac{\gamma_C p_1^*}{m_1 \gamma_1 + m_2 \gamma_2} \left( \frac{m_1 \gamma_1^* m_2 \gamma_2^*}{p_1^{*2}} c^2 + 1 \right)^2$$

- This model also breaks down for very low speeds/temperatures - the collisionality diverges to infinity
- Equation 6 gives a mean free path equal to the particle spacing; by using this as a lower bound the divergence is avoided

$$(6) \quad s' = \left( \frac{4\pi}{3} \right)^{1/3} \frac{n_1 n_2}{n_{12}} \bar{\Delta} t \frac{m_1 + m_2}{\max \left( m_1 n_1^{2/3}, m_2 n_2^{2/3} \right)} v_{rel} \quad s = \min(s, s')$$

# Collision step in main iteration loop

## os-simulation

```
iter_sim()
....
call ionize_neutral(...)
call sort_collide( sim%part, dx, n, t,
dt, n_threads)
call sim%jey%update_boundary(...)
....
```

## os-particles

```
if
  collide()
else
  sort()
...

```

## os-spec-collisions

```
collide_particles()
collision_sort()
!$omp parallel do
do loop_over_cells
  find_fluid_state()
  debye_length()
  random_paring()
  ★ select case (model)
do loop_over_species()
  ★ unlike_collide()
  like_collide()
enddo
```

OpenMP parallelization for all collision cells

# Collision loop in each cell



```
unlike_collide_perez()  
  <find correction for uneven particle weights>  
  <find (weighted) average relative velocity>  
  log_lambda = ...  
  s_first = ... !shared part of collision frequency  
  do loop_over_pairs  
    relative_velocity = ... ! one-particle-at-rest frame  
    <boost momenta to center-of-momentum frame>  
    s = ... ! collision frequency for pair  
    <invert Nanbu equation to get collision angle>  
    <calculate rotation in CoM>  
    <update momenta stochastically based on weighting>  
    <boost momenta back to lab frame>
```

# Five input parameters are required

- Needed input parameters:
  - Actual charge density for  $n=1$  (cgs)
  - Which species to collide
    - Self-collide?
  - Charge state of ion
  - Collision timestep (0, the default, doubles as flag to turn collisions off)

```
1 simulation
2 {
3     n0 = 1.1d21,
4 }
5
6 species
7 {
8     if_collide = .true.,
9     if_like_collide = .false.,
10    q_real = 1,
11 }
12
13 collisions
14 {
15     n_collide = 1,
16     ! Defaults
17     collision_model = "Perez",
18     nx_collision_cells(1:2) = 1, 1,
19     coulomb_logarithm_automatic = .true.,
20     ! Optional (not default)
21     perez_low_temp_correction = .true.,
22 }
23
```



# Other inputs to know



- Defaults to be aware of
  - Perez collision model
  - Collision cell size equals PIC cell size
  - Calculate Coulomb log in each cell
- One option one might want: Low temperature correction

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22 }
23
```

# Things you have to worry about a little



- Numerical heating (see Paulo Alves)
- Ratio of collision cell to MPI node
- Good resolution
  - Depends on the problem, no hard rules
  - Size of collision gradients (`nx_collision_cells`)
  - Resolving the collision frequency (`n_collide`)
  - Representing the distribution function (`nx_collision_cells`, `ppc`)
  - That said, the algorithm is fairly forgiving

# Things you don't have to worry about



- Over-resolving the collision frequency
- Relative size of macro-particles
- Multiple equivalent species
- Highly different densities
- Relativistic effects
- Conservation of energy and momentum \*on average\*  
(except for heating)

# Computational Cost



- It is fairly expensive
- Depends on system, but roughly 2X cost for two species colliding at each timestep
- Currently only vectorization is OpenMP (which works quite well!); adopting for GPGPU or KNL seems promising, but future work
- SIMD within each collision cell is possible too, but is challenging

# Take Away Message



- For the general user, just a few parameters need to be input and the collision package will take care of the rest. Be careful of numerical heating, but otherwise you can't really mess up.
- With a little more care, you may improve the computational performance without losing physics.
- For the developer, both the package and the individual models are well isolated; easy (and encouraged) to add to or modify.
- Better vectorization, impact ionization are obvious goals